Viscous Flow Lecture 12 Slow flow past a sphere Axisymmetre flow u(r,o) = urer+ uo lo and p(r,0) in sphenical polar coordinates. 1=0 on r=1 TTTT QS T > D Donensvonless variables

U > k = er cos 0 - eo sur o and Re ($u \cdot \nabla$) $u = -\nabla p + \nabla^2 u$ V. u = 0 Expand $u = u^{(0)} + Re u^{(1)} + ...$ $P = P^{(0)} + Re P^{(1)} + \cdots$ Keep only the leading order terms u(0) and p(0) and drop "(0)" superscripts. Slow (Stokes) flow equations

 $V \cdot u = 0$ and $\nabla p = \nabla' u$

In spherical polar coordinates, incompressibility becomes Tuesday, 17 November 2020 = \frac{1}{75\text{cho}} \left[\frac{3}{77} \left(Urr^2\sin\theta\right) \left(\text{den concolled}\right) \\ + \frac{1}{70} \left(U\text{or}^2\sin\theta\right)\right] 0 = V. L where $r^2 sch \theta$ is the Jacobian for $(x, y, z) \mapsto (r, \theta, \phi)$ coordinates Written Whe this as ond too come from V en sphenzal polas. We can satisfy incompressibility by introducing a Stokes stream function $\psi(r, 0)$ s.t. $Ur = \frac{1}{r^2 \sin \theta} \frac{\partial \Psi}{\partial \sigma}, \quad U_{\theta} = -\frac{1}{r \sin \theta} \frac{\partial \Psi}{\partial r}$ er reo romo ex or do de que o o o que o o U = rsuno = V1 (rsuno eq) This is the sphenzel analogue of $u = P \Lambda (Y k)$ in ZD. round ell is a longth on a latitude circle from varying of To elemenate the pressure ne'll want the vorticity equation. $W = \sqrt{\Lambda} U = \frac{1}{r \sin \theta} \begin{vmatrix} er & reo & r \sin \theta e_f \\ \partial r & \partial \theta & \partial \phi \end{vmatrix}$ $Ur \quad r u_{\theta} \quad r \sin \theta u_{\theta}$ =-ed - 1 (24 - coto 24)
75010 (272 - rz 20) $+\frac{1}{r^2}\frac{\partial^2 4}{\partial \theta^2}$ = - ea los D24 This defines the Stohes operator D. Note that $D^2 + \nabla^2$ because D^2 3 applied to rector fields on the Co devection, and ex is itself a furction of D and Q. The slow flow equetions are $\nabla P = \nabla^2 u = \nabla (\nabla \cdot u) - \nabla \Lambda (\nabla \Lambda u)$ = - curl u as Vou=0 Taking the circl again to eliminate the pressure gives 0 = curl (Pp) = - curl's 4 = - curl 4 (TSURD Pd) $\Rightarrow D^4 \Psi = O \qquad \left(D^4 = \left(D^2\right)^2\right)$ In 20 Cortescons we had P44=0 Boundary conditions: u = 0 on $r = 1 \Rightarrow \Psi = \frac{3\Psi}{3r} = 0$ The far field condition for uniform on coming flow is u > le = er coso - la suro as r=20, 50 $U_r = \frac{1}{r^2 \sin \theta} \frac{\partial \Psi}{\partial \theta} \rightarrow \cos \theta$ $u_{o} = -\frac{1}{rsuno} \frac{34}{3r} \Rightarrow -sun o_{3}$ giving $4n \pm r^3 sin^3 0$ as $r > \infty$ This behaviour in the for hild suggests the separable solution $\Psi(r,o) = f(r) sun^2 0$ $D^2 \Psi = \left(\frac{2}{2m} - \frac{2}{r^2}\right) f(r) \sin^2 \theta$ $D^{4} \Psi = \left(\frac{3^{2}}{5r^{2}} - \frac{2}{r^{2}}\right)^{2} f(r) \sin^{2} \theta$ This is homogeneous in r, so by $f(r) = r^n$ for solutions. This gives r4, r2, r, 1/r Applying the boundary conditions f = f' = 0 on r = 1f(r) $n = r^2$ des $r \to \infty$ gives $\psi(r_0 \theta) = (\frac{1}{2}r^3 - \frac{3}{4}r + \frac{1}{4r})sh^3\theta.$ The straightforward solution works at leading order for flow around a sphere. A paradox arises if ne by to find the O(Re) convections $U^{(i)}$ and $p^{(i)}$ since 4, has an r'smacos à dependence that con't match to He crifonn flow at infikity. Again, we need a resceling for a bounday læger æt chfriety. This version is called the Whitehead paradox.

The Stokes drag formula (1851) Tuesday, 17 November 2020 Involved in at least 3 Nobel Prize - winning research projects. 4= (=12-3/4r+4r) sun 8 The velocity components are: $Ur = \frac{1}{r^2 \sin \theta} \frac{34}{30}$ $= \left(1 - \frac{3}{2r} + \frac{1}{2r^3}\right) \cos \theta$ $u_0 = -\frac{1}{rsuno} \frac{\partial u}{\partial r}$ = (-1+ \frac{3}{4r} + \frac{1}{4r^3}) sin \O $NP = - cwl^2 u$ we get 3 cos 8 P = Poo - Zrz with postle arbitrary constant pressure at infinity. $P-Po = O(1/r^2),$ nith Us = k 14-k = 0 (/r) The drag force is in the ke direction by symmetry, so equal to Dk where n outward normal D= Jj k. s.n ds To For k = er cos 8 - eo suo D = II orr cos 0 - ord sun o as $\sqrt{u} = -b + \frac{2u}{2u}$ where $\frac{\sigma_{rr}}{r=1} = -\frac{2}{r}\cos\theta$ $Ord | = -\frac{3}{2} sud$ D = South of the South of Orr coso-Oroscho) $= -Poc \cos a$ $+ \frac{3}{2}$ The viscous contribution is constant, and lle pressure contribution integrates to zero. Avea is 4TT $D = 4\pi \times \frac{3}{2} = 6\pi$ Restoring the dimensions gives Stokes) famous drag force $D = 6\pi \left(\frac{\mu U}{\alpha}\right) a^2 = 6\pi \mu a U$ scale scale for over for I and p For example, the terminal velocity of a solid sphere of criforn density es falling though viscous flevid with density PL is given 6 mma 0 = \frac{4}{3} \pi \text{Tra}^3 (P_S - P_L) 9 $U = \frac{z}{q} \frac{q^2 (R_s - R_L) g}{2}.$ This can be used to measure VIS cosèties. This can also be used to measure the electric charge on a drop of oil on air o drop roltege Measoning the falling speed with and without the electric field determines the vadius and the chage. Millikan & Fletcher measured lle charge on the electron Whe this to get (1913) $e = 1.5924(17) \times 10^{-19} C$ The modern figure is e = 1-6 622--- x 10-17 C. The Stokes dag formula was also used in Einstein's Heavy of Brownian motion. Verified experimentally by Penn.