

# (Classical) Relativistic Point Particle

$$S = -m \int_Y ds$$

Treat as "0+1-dimensional field theory",  $\tau$  local time parameter on worldline.  $x^M(\tau): \mathbb{R}_\tau \rightarrow \mathbb{R}^d$  worldline fields.

$$S[x^M(\tau)] = -m \int_{\tau_i}^{\tau_f} \sqrt{-g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau}} = -m \int_{\tau_i}^{\tau_f} d\tau \sqrt{-h}$$

$\uparrow$  induced (pull back) metric on  $\mathbb{R}_\tau$

Euler - Lagrange equations  $\rightarrow$  trajectories are geodesics.

## Symmetries of the action ...

- if  $g_{\mu\nu} = \eta_{\mu\nu}$  (flat Minkowski metric) then spacetime Poincaré invariance

$$x^M(\tau) \rightarrow \Lambda^M_\nu x^\nu(\tau) + b^M \quad \text{for } \Lambda \in SO(1, d-1) \\ b \in \mathbb{R}$$

more generally, spacetime isometry group is realized as "internal" symmetries on the worldline theory.

- 1-dim! reparameterization invariance

$$\tau \rightarrow \tilde{\tau}(\tau) \\ x^M(\tau) \rightarrow \tilde{x}^M(\tilde{\tau}) = x^M(\tau)$$

this is because action is really a function of worldline as a subset of spacetime; don't care about parameterization.

THIS IS A GAUGE SYMMETRY / REDUNDANCY OF DESCRIPTION.

This is a fine classical action, and can be analyzed as usual. Not very appealing if we want to quantize.

# (Classical) Relativistic Point Particle

Introduce "einbein"  $e: \mathbb{R}_\tau \rightarrow \mathbb{R}_{>0}$ , a "square root of worldline metric", with new action:

$$\tilde{S} = \frac{1}{2} \int (e^{-1} g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} - em^2) d\tau$$

Imposing equations of motion for  $e$  gives  $e^2 = -\frac{1}{m^2} g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau}$ . Substitute back into  $\tilde{S}$  to get  $S$ , so classically equivalent.

Symmetries of  $\tilde{S} \dots$

- spacetime Poincaré/isometries ( $e(\tau)$  invariant)
- reparameterization of worldline

|                                                                                 |                                                             |
|---------------------------------------------------------------------------------|-------------------------------------------------------------|
| $\tau \mapsto \tilde{\tau}(\tau)$                                               | $\tau \mapsto \tau - \xi(\tau)$ for $\xi$ infinitesimal     |
| $x^\mu(\tau) \mapsto \tilde{x}^\mu(\tilde{\tau}) = x^\mu(\tau)$                 | $x^\mu \mapsto x^\mu + \xi \frac{dx^\mu}{d\tau}$            |
| $e(\tau) \mapsto \tilde{e}(\tilde{\tau}) = \frac{d\tau}{d\tilde{\tau}} e(\tau)$ | $e \mapsto e + \xi \frac{de}{d\tau} + \frac{d\xi}{d\tau} e$ |

Can use reparameterization to gauge fix and set  $e(\tau) \equiv 1/m$ .

$$\tilde{S}_{\text{fixed}} = \frac{1}{2} \int (m g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} - m) d\tau$$

Equations of motion give geodesic equation:

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma^\mu_{\alpha\beta} \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0$$

Now there is an additional constraint from the einbein equation of motion

$$g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} + 1 = 0 \quad (\text{so } \frac{dx^\mu}{d\tau} \text{ is 4-velocity; } \tau \text{ is proper time}).$$

Alternatively, if  $m=0$  can gauge fix  $e=1$  and recover geodesic equation plus null constraint

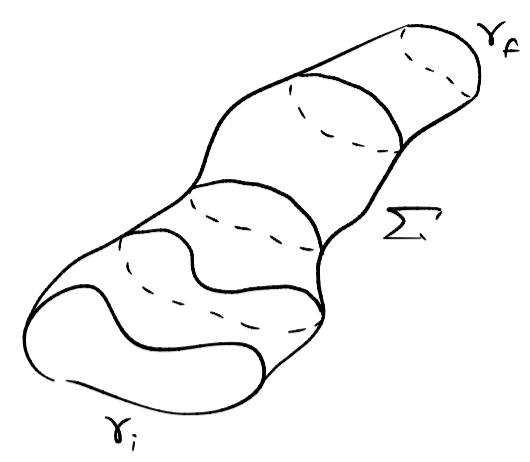
$$g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} = 0 \quad (\text{so null geodesic}).$$

This could be a reasonable starting point for quantization (quadratic in time derivatives, etc.). Could even imagine starting to add interactions, building up Feynman diagrams in a first-quantized setting.

# (Classical) Relativistic String

Analogue of first action

$$S[\Sigma] = -T \int_{\Sigma} dA$$



Introduce worldsheet parameters  $(\sigma, \tau)$ , treat as 1+1-dimensional field theory with fields  $X^M(\tau, \sigma): \Sigma \rightarrow \mathbb{R}$

$$S_{NG}[X^M(\tau, \sigma)] = -T \int_{\Sigma} \sqrt{-(\partial_z X \cdot \partial_z X)(\partial_\sigma X \cdot \partial_\sigma X) + (\partial_z X \cdot \partial_\sigma X)^2} = -T \int_{\Sigma} \sqrt{-h} dz d\sigma$$

Nambu-Goto action

$$V \cdot W = g_{\mu\nu} V^\mu W^\nu \text{ for spacetime vectors}$$

Euler-Lagrange equations  $\rightarrow$  minimal area surfaces.

T interpreted as string tension

$$\left\{ \begin{array}{l} \text{Set } X^0 = \tau, \partial_\tau X^{M \neq 0} = 0, g_{0\mu} = 0 \text{ for } \mu \neq 0, \text{ so static string in static geometry} \\ S = -T \int dz d\sigma \sqrt{\partial_\sigma X \cdot \partial_\sigma X} = -T \int dz (\text{string length}) \\ = \int dz (\text{Kinetic-Potential}), \text{ so potential energy} \sim T \times \text{length} \end{array} \right.$$

Symmetries of the Nambu-Goto action:

- spacetime Poincaré/isometries as internal symmetries on worldsheet

$$X^M(\tau, \sigma) \mapsto \Lambda^M_\nu X^\nu(\tau, \sigma) + b^M$$

- 2-dim'l reparameterization invariance (often referred to as diffeomorphism invariance)

$$(\tau, \sigma) \rightarrow (\tilde{\tau}(\tau, \sigma), \tilde{\sigma}(\tau, \sigma))$$

$$X^M(\tau, \sigma) \rightarrow \tilde{X}^M(\tilde{\tau}, \tilde{\sigma}) = X^M(\tau, \sigma)$$

Again, reparameterizations are a gauge symmetry.

Once again, a fine classical theory (see problem sheet 1). Not clear how to quantize.

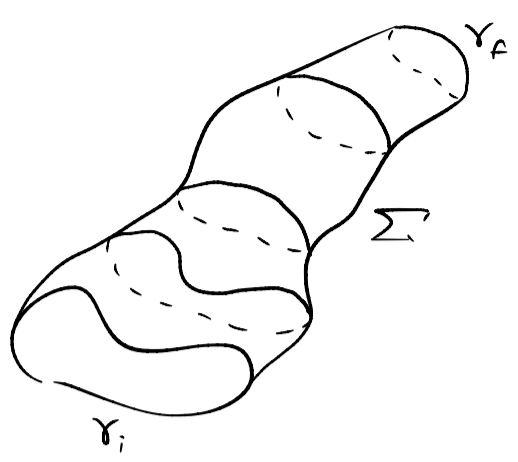
# (Classical) Relativistic String

Analogue of second action

$$S_P[\gamma_{ab}, X^M] = -\frac{T}{2} \int_{\Sigma} d\tau d\sigma \sqrt{-\gamma} \gamma^{ab} \frac{\partial X^M}{\partial \xi^a} \frac{\partial X^M}{\partial \xi^b}$$

$\gamma_{ab}$ : auxiliary Lorentzian world-sheet metric  
 $\gamma^{ab}$ :  $\det(\gamma_{ab})$

Polyakov Action (Brink-di Vecchia-Haue-Deser-Zumino)



As in point particle case, prove classical equivalence by solving E.O.M. for worldsheet metric  $\gamma_{ab}$ .

Symmetries of the Polyakov action

- spacetime Poincaré ( $\gamma$  doesn't transform)
- worldsheet reparameterizations

$$\gamma_{ab}(\xi) \rightarrow \tilde{\gamma}_{ab}(\tilde{\xi}) = \gamma_{cd}(\xi) \frac{\partial \xi^c}{\partial \tilde{\xi}^a} \frac{\partial \xi^d}{\partial \tilde{\xi}^b} \quad (\xi = (\sigma, \tau))$$

- Weyl invariance (local scale transformations).

$$\gamma_{ab} \mapsto e^{2\omega(\sigma, \tau)} \gamma_{ab}$$

special to 2-dimensional worldvolume;  $\sqrt{|\gamma|}$  picks up factor of  $\exp(D\omega)$  while  $\gamma^{ab}$  picks up  $e^{-2\omega}$ .

# worldvolume dimensions (D=2 for string)

Weyl transformations are now also a gauge symmetry, so  $D=2$  has extra gauge invariance. This is, in a way, what is special about strings (vs. membranes, extended objects of greater dimension, etc.).

Should consider what power-counting renormalizable terms we could add to the action consistent with symmetries.

$$S^{(1)} = \lambda \int_{\Sigma} \sqrt{-\gamma} d\tau d\sigma \quad \text{"cosmological constant" term} \rightarrow \text{not Weyl invariant.}$$

$$S^{(2)} = \frac{1}{2\pi} \int_{\Sigma} \sqrt{-\gamma} R^{(2)}(\gamma) d\tau d\sigma \quad \text{"Einstein-Hilbert term"} \rightarrow \text{total derivative in 2d, no effect locally.}$$

$R^{(2)}$ : worldsheet Ricci scalar  
 (2d gravity is "trivial")

A note on dimensional analysis: worldsheet & spacetime both have notion of length scale/units.

|                                                                                                            |                                                                                                         |
|------------------------------------------------------------------------------------------------------------|---------------------------------------------------------------------------------------------------------|
| worldsheet: $[X^M] = 0$<br>$[\frac{d}{d\xi}] = 1$<br>$[\gamma] = 0$<br>$[T] = 0$<br>$[d\tau d\sigma] = -2$ | spacetime: $[X^M] = -1$<br>$[\frac{d}{ds}] = 0$<br>$[\gamma] = 0$<br>$[T] = 2$<br>$[d\tau d\sigma] = 0$ |
|------------------------------------------------------------------------------------------------------------|---------------------------------------------------------------------------------------------------------|

$\alpha'$  will play role of  $\hbar$  on the world sheet

Conventional to write  $T = \frac{1}{2\pi\alpha'}$ ,  $[\alpha'] = -2$

# Gauge Fixing the Polyakov String

Use reparameterization & Weyl gauge invariances to simplify the action further.

Reparam.:  $\gamma^{ab} \rightarrow e^{2\omega(\tau,\sigma)} \gamma^{ab}$  "conformal gauge"  $\left\{ \begin{array}{l} \text{naive counting: 2 functions of} \\ \text{coordinates in reparam., setting} \\ \text{two metric entries to zero} \end{array} \right\}$

$\left( \begin{smallmatrix} 1 & 0 \\ 0 & 1 \end{smallmatrix} \right)$

Weyl:  $e^{2\omega(\tau,\sigma)} \gamma^{ab} \rightarrow \gamma^{ab}$  "unit gauge"  $\left\{ \begin{array}{l} \text{one more function to cancel, this} \\ \text{one clearly works (not naive).} \end{array} \right\}$

Comments: room for some concern here; can we do these gauge fixings globally on  $\Sigma$  or just locally? For  $\gamma^{ab}$  Lorentzian/Semi-Riemannian, things could get pretty messy.

The secret is that ultimately when the global structure of  $\Sigma$  is relevant, we will "Wick rotate" to Euclidean signature on the worldsheet. Here the story is cleaner (though still interesting). For now we press on.

The Polyakov action in conformal gauge simplifies dramatically

$$S_P^{\text{conf. gauge}}[X^\mu] = -\frac{T}{2} \int d\tau d\sigma \left( -\partial_\tau X \cdot \partial_\tau X + \partial_\sigma X \cdot \partial_\sigma X \right)$$

This is just the theory of  $D$  massless scalar fields in flat  $1+1$ -dimensional space. (One with wrong-sign kinetic term)

In addition, supplement with constraints from equations of motion for worldsheet metric.

$$T_{ab} := -\frac{2}{T} \cdot \frac{1}{\sqrt{-\gamma}} \frac{\delta S_P}{\delta \gamma^{ab}} \stackrel{!}{=} 0$$

world-sheet stress tensor E.O.M. constraint

$$T_{ab} = \partial_a X \cdot \partial_b X - \frac{1}{2} \gamma_{ab} \gamma^{cd} \partial_c X \cdot \partial_d X \quad \text{in conformal gauge.}$$

$$\left. \begin{array}{l} T_{\tau\tau} = \frac{1}{2} \partial_\tau X \cdot \partial_\tau X + \frac{1}{2} \partial_\sigma X \cdot \partial_\sigma X \\ T_{\tau\sigma} = \partial_\tau X \cdot \partial_\sigma X \\ T_{\sigma\sigma} = \frac{1}{2} \partial_\tau X \cdot \partial_\tau X + \frac{1}{2} \partial_\sigma X \cdot \partial_\sigma X \end{array} \right\} \gamma^{ab} T_{ab} = T_{\sigma\sigma} - T_{\tau\tau} = 0 \quad \left( \text{this is due to Weyl invariance} \right)$$

So only two constraints to impose, with third satisfied identically due to Weyl symmetry.

# Gauge Fixing the Nambu-Goto String

Note for good measure, can use reparameterizations to fix Nambu-Goto action to conformal gauge as well.

Induced metric  $h_{ab} = \partial_a X \cdot \partial_b X$

Fix  $h_{\tau\sigma} = \partial_\tau X \cdot \partial_\sigma X = 0$

$h_{\tau\tau} + h_{\sigma\sigma} = \partial_\tau X \cdot \partial_\tau X + \partial_\sigma X \cdot \partial_\sigma X = 0$

Use constraints to rewrite action

$$S_{NG} = -T \int \sqrt{-h} = -T \int_{\Sigma} \sqrt{(\partial_\sigma X \cdot \partial_\tau X)^2 - (\partial_\tau X \cdot \partial_\tau X)(\partial_\sigma X \cdot \partial_\sigma X)} d\tau d\sigma$$

$$= -T \int_{\Sigma} \sqrt{0 - \frac{1}{2}(\partial_\tau X \cdot \partial_\tau X - \partial_\sigma X \cdot \partial_\sigma X) - \frac{1}{2}(\partial_\sigma X \cdot \partial_\sigma X - \partial_\tau X \cdot \partial_\tau X)} d\tau d\sigma$$

$$= -\frac{T}{2} \int_{\Sigma} d\tau d\sigma (-\partial_\tau X \cdot \partial_\tau X + \partial_\sigma X \cdot \partial_\sigma X)$$

So subject to some constraints, recover same simple free field action.

(More on NG action on problem sheet 1)

- Plan for next week:
- General classical solutions
  - Constraints
  - Residual gauge symmetry & conformal transformations
  - Quantization