

## Problem Sheet 4: Solutions

## Question 1. Rotating sphere in Stokes flow.

With  $\hat{\mathbf{x}} = \mathbf{x} - \mathbf{x}_0$ ,  $r = |\mathbf{x} - \mathbf{x}_0|$  and

$$G_{ij} = \frac{\delta_{ij}}{r} + \frac{\hat{x}_i \hat{x}_j}{r^3},$$

the rotational dipole  $\mathbf{G}^c$  is defined by

$$G_{im}^c := \frac{1}{2} \epsilon_{mlj} \frac{\partial G_{ij}}{\partial x_{0,l}},$$

where

$$\epsilon_{mlj} := \begin{cases} +1 & \text{if } (m, l, j) = (1, 2, 3) \text{ or } (3, 1, 2) \text{ or } (2, 3, 1) \\ -1 & \text{if } (m, l, j) = (1, 3, 2) \text{ or } (2, 1, 3) \text{ or } (3, 2, 1) \\ 0 & \text{if any of } i, j, k \text{ are equal} \end{cases}$$

**Part a.** We have

$$\begin{aligned} \frac{1}{2} \epsilon_{mlj} \frac{\partial}{\partial x_{0,l}} G_{ij} &= \frac{1}{2} \epsilon_{mlj} \frac{\partial}{\partial x_{0,l}} \left[ \frac{\delta_{ij}}{r} + \frac{\hat{x}_i \hat{x}_j}{r^3} \right], \\ &= \frac{1}{2} \epsilon_{mlj} \left[ \delta_{ij} \frac{\partial}{\partial x_{0,l}} \left( \frac{1}{r} \right) - \frac{1}{r^3} (\delta_{ij} \hat{x}_j + \delta_{jl} \hat{x}_i) + \hat{x}_i \hat{x}_j \frac{\partial}{\partial x_{0,l}} \left( \frac{1}{r^3} \right) \right], \\ &= \epsilon_{mli} \frac{\hat{x}_l}{r^3}. \end{aligned}$$

**Part b.** Thus  $G_{im}^c q_m$  is a solution of Stokes equations for any constant vector  $\mathbf{q}$  and it decays at spatial infinity. With  $\mathbf{x}_0$  the centre of the sphere, the sphere is given by  $r = a$ , where we have

$$a^3 G_{im}^c \Omega_m = a^3 \epsilon_{iml} \Omega_m \frac{\hat{x}_l}{r^3} = \epsilon_{iml} \Omega_m \hat{x}_l,$$

which is the velocity on a rotating sphere.

Hence

$$v_i := a^3 G_{im}^c \Omega_m$$

in a solution of the Stokes flow for a rotating sphere with radius  $a$  and angular velocity  $\boldsymbol{\Omega}$  and, by uniqueness, it is the solution (with constant pressure).

**Part c.** One can use brute force, though that would be on the long side. The stress field of the Stokes solution

$$u_i = \frac{1}{8\pi\mu} G_{ij} g_j$$

be given by  $\sigma_{ij} = T_{ijp}g_p$ , with symmetry in the indices  $i, j$ . You can show

$$\frac{\partial T_{ijs}}{\partial x_j} = -\delta_{is}\delta(\hat{\mathbf{x}}),$$

from the momentum balance for  $u_i$ , consistent with the notion that  $u_i$  is the flow associated with a point force.

Then the stress field associated with  $\mathbf{v} = a^3 \mathbf{G}^c \cdot \boldsymbol{\Omega}$ , that is

$$v_m = a^3 G_{mn}^c \Omega_n,$$

is given by

$$\Sigma_{ij} = (8\pi\mu)a^3 T_{ijp}^c \Omega_p = (8\pi\mu) \frac{a^3}{2} \epsilon_{plq} \frac{\partial}{\partial x_{0,l}} T_{ijq} \Omega_p = -(8\pi\mu) \left( \frac{a^3}{2} \Omega_p \right) \epsilon_{plq} \frac{\partial}{\partial x_l} T_{ijq}.$$

Hence the  $s^{th}$  component of the moment of the sphere due to the fluid is given by

$$M_s = \int_{Sphere} \epsilon_{sri} x_r \Sigma_{ij} n_j dS = -(8\pi\mu) \left( \frac{a^3}{2} \Omega_p \right) \int_{Sphere} \epsilon_{sri} x_r \epsilon_{plq} \left( \frac{\partial}{\partial x_l} T_{ijq} \right) n_j dS.$$

Using the divergence theorem we have

$$M_s = -(8\pi\mu) \left( \frac{a^3}{2} \Omega_p \right) \int_{Sphere} \epsilon_{sji} \epsilon_{plq} \left( \frac{\partial}{\partial x_l} T_{ijq} \right) dV = -(8\pi\mu) \left( \frac{a^3}{2} \Omega_p \right) \int_{Sphere} \epsilon_{sri} x_r \epsilon_{plq} \frac{\partial}{\partial x_j} \left( \frac{\partial}{\partial x_l} T_{ijq} \right) dV.$$

The first term is zero as there is an contraction between  $\epsilon$  with antisymmetry in  $i, j$  and  $T_{ijq}$ , with symmetry in  $i, j$ . Commuting derivatives and noting the above relation for the derivative of  $T_{ijk}$  we have

$$M_s = (8\pi\mu) \left( \frac{a^3}{2} \Omega_p \right) \int_{Sphere} \epsilon_{sri} x_r \epsilon_{pli} \frac{\partial}{\partial x_l} (\delta(\hat{\mathbf{x}})) dV = -(8\pi\mu) \left( \frac{a^3}{2} \Omega_p \right) \int_{Sphere} \epsilon_{sri} \epsilon_{pri} \delta(\hat{\mathbf{x}}) dV,$$

noting the additional surface integral term to deduce the final equality must be zero as the  $\delta$ -function has no support on the surface of the sphere. Finally, with  $\epsilon_{sri} \epsilon_{pri} = 2\delta_{ps}$ , we have

$$M_s = -8\pi\mu a^3 \Omega_s,$$

as required.

## Question 2. Ciliary Pumping.

Detailed calculation is not necessary to deduce the expression for  $U$ . Fourier modes decouple at the first non-trivial order and each derivative acting on mode number  $n$  just induces a factor of  $n$ . Thus one can determine the contribution to  $U$  from the Fourier mode

$$x_e - x = \epsilon(-b_n \cos(n[x + t])), \quad y_e = \epsilon c_n \sin(n[x + t]) \quad (1)$$

by identifying

$$a = -nb_n, \quad b = nc_n \quad (2)$$

in the example sheet result

$$U_2 = \frac{1}{2} (b^2 + 2ab - a^2)$$

to obtain the contribution from this mode.

We then can consider the remaining Fourier modes without detailed calculation. The mode

$$x_e - x = \epsilon a_n \sin(n[x + t]), \quad y_e = \epsilon(-d_n \cos(n[x + t]))$$

is simply a phase shift of the mode in equation (1). By considering a shifted time coordinate

$$t = \bar{t} + \frac{1}{n} \frac{\pi}{2}$$

we can determine the contribution from this mode by the substitution

$$d_n \rightarrow c_n, \quad a_n \rightarrow -b_n.$$

followed by the identification (2). Hence we use

$$a = na_n, \quad b = nd_n$$

in the example sheet result

$$U_2 = \frac{1}{2} (b^2 + 2ab - a^2)$$

to obtain the contribution from this mode.

Summing all contributions, and noting  $U = \epsilon^2 U_2$  to leading order, gives

$$U = \frac{1}{2} \epsilon^2 \sum_{n=1}^{\infty} n^2 [c_n^2 + d_n^2 - a_n^2 - b_n^2 + 2(a_n d_n - c_n b_n)].$$

We now determine power optimal strokes, defined as those maximising absolute velocity, subject to the constraint of a fixed power consumption  $W$  using Lagrange multipliers with the above leading order expressions. Thus we consider

$$L[\{a_n, b_n, c_n, d_n\}] = U[\{a_n, b_n, c_n, d_n\}] - \lambda(P[\{a_n, b_n, c_n, d_n\}] - W) \quad (3)$$

and the extremal conditions

$$\frac{\partial L}{\partial a_n} = \frac{\partial L}{\partial b_n} = \frac{\partial L}{\partial c_n} = \frac{\partial L}{\partial d_n} = 0. \quad (4)$$

Thus

$$a_n^2 + 2a_nd_n - d_n^2 = 0, \quad b_n^2 - 2b_nc_n - c_n^2 = 0.$$

and hence  $a_n = (-1 \pm \sqrt{2})d_n$  and  $b_n = (1 \pm \sqrt{2})c_n$ , which yields

$$U = \epsilon^2 \sum_{n=1}^{\infty} n^2 \left( (2 \pm 2\sqrt{2})c_n^2 + (2 \mp 2\sqrt{2})d_n^2 \right) \quad (5)$$

$$P = \epsilon^2 \sum_{n=1}^{\infty} n^3 \left( (4 \pm 2\sqrt{2})c_n^2 + (4 \mp 2\sqrt{2})d_n^2 \right). \quad (6)$$

Therefore the optimal stroke is achieved when  $a_n = b_n = c_n = d_n = 0$  for  $n \geq 2$ . Without loss of generality, we can set  $b_1 = 0$  as it is just a phase difference, and we finally have the optimal strokes  $a_1 = (-1 \pm \sqrt{2})d_1$ .

As an extra to the question, note the extremal velocity is

$$U = \mp \frac{W}{\sqrt{2}}.$$

### Question 3. Ciliate Motility.

Take a reference frame comoving with the swimmer oriented such that the direction of the swimmer velocity is given by  $\mathbf{U} = U\mathbf{e}_z$ . The non-dimensional Stokes equations are

$$\nabla^2 \mathbf{u} = \nabla p, \quad \nabla \cdot \mathbf{u} = 0,$$

with

$$\mathbf{u} = -U\mathbf{e}_z, \quad \text{as } |\mathbf{x}| \rightarrow \infty, \quad \mathbf{u} = \epsilon\beta(t) \sin \theta \mathbf{e}_\theta, \quad \text{on } r = 1,$$

where  $\mathbf{e}_\theta$  is the unit vector in the direction of increasing spherical polar  $\theta$ , where  $r = |\mathbf{x}|$  and  $z = r \cos \theta$ ,  $x = r \sin \theta \cos \varphi$  for instance.

By symmetry we have  $\mathbf{U}$  is in the  $\mathbf{e}_z$  direction.

Show that

$$\mathbf{u} = \left[ -U(t) + \frac{Q(t)}{r^3} \right] \cos \theta \mathbf{e}_r + \left[ U(t) + \frac{P(t)}{r^3} \right] \sin \theta \mathbf{e}_\theta, \quad p = \text{Const}$$

is a solution of the Stokes equation for  $Q(t) = 2P(t)$ . To do this, you will need to consider the vector Laplacian of  $\mathbf{u}$ . This is non-trivial in non-Cartesian coordinates and you may wish to consider using a symbolic algebra package such as Mathematica:

$$\text{f1}[r\_,\text{theta}\_,\text{phi}\_] := (-U + Q/r^3) * \text{Cos}[\text{theta}]$$

$$\text{f2}[r\_,\text{theta}\_,\text{phi}\_] := (U + P/r^3) * \text{Sin}[\text{theta}]$$

$$\text{FullSimplify}[\text{Laplacian}[\text{f1}[r,\text{theta},\text{phi}], \{r,\text{theta},\text{phi}\}, \text{"Spherical"}] - 2 * \text{f1}[r,\text{theta},\text{phi}]/r^2 - 2/(r^2 * \text{Sin}[\text{theta}]) * (\text{f2}[r,\text{theta},\text{phi}] * \text{Cos}[\text{theta}] + \text{Sin}[\text{theta}] * D[\text{f2}[r,\text{theta},\text{phi}], \text{theta}])]$$

$$\frac{2(-2P + Q)\text{Cos}[\text{theta}]}{r^5}$$

$$\text{FullSimplify}[\text{Laplacian}[\text{f2}[r,\text{theta},\text{phi}], \{r,\text{theta},\text{phi}\}, \text{"Spherical"}] + 2 * D[\text{f1}[r,\text{theta},\text{phi}], \text{theta}]/r^2 - 1/(r^2 * \text{Sin}[\text{theta}]) * \text{Sin}[\text{theta}] * \text{f2}[r,\text{theta},\text{phi}]]$$

$$\frac{2(2P - Q)\text{Sin}[\text{theta}]}{r^5}$$

Given  $Q(t) = 2P(t)$ , the above two expressions are zero and the governing equation is satisfied. One can then read off that  $U(t) = 2\epsilon\beta/3$  from the boundary condition, by finding  $Q(t)$  in terms of  $U(t)$  from setting the coefficient of  $\mathbf{e}_r$  to zero and then equating the coefficient of  $\mathbf{e}_\theta$  to the boundary conditions at the boundary.

## Question 4. Resistive force theory.

For the force balance with a spherical cell body of radius  $a$ , we have

$$\mathbf{0} = (\text{Drag force on body}) + (\text{Drag force on flagellum}). \quad (7)$$

and  $\mathbf{e}_t = (-1, \epsilon h_s)$ ,  $\mathbf{e}_n = (\epsilon h_s, 1)$  and the velocity of the flagellum element is given by  $\mathbf{U} = (U, V + \epsilon h_t)$ .

Hence the drag force per unit length on the element  $ds$  is given by

$$\begin{aligned} \mathbf{f} &= -[C_N \mathbf{e}_n \cdot \mathbf{U} \mathbf{e}_n + C_T \mathbf{e}_t \cdot \mathbf{U} \mathbf{e}_t] = -[(C_N - C_T) \mathbf{e}_n \cdot \mathbf{U} \mathbf{e}_n + C_T \mathbf{U}] \\ &= -[(C_N - C_T) \mathbf{e}_n \otimes \mathbf{e}_n + C_T \mathbf{I}] \mathbf{U} \\ &= -\left[(C_N - C_T) \begin{pmatrix} \epsilon^2 h_s^2 & \epsilon h_s \\ \epsilon h_s & 1 \end{pmatrix} + C_T \mathbf{I}\right] \begin{pmatrix} U \\ \epsilon h_t + V \end{pmatrix} \\ &= -(C_N - C_T) \begin{pmatrix} \epsilon^2 h_s^2 U + \epsilon^2 h_s h_t + \epsilon h_s V \\ \epsilon h_s U + \epsilon h_t + V \end{pmatrix} - C_T \begin{pmatrix} U \\ \epsilon h_t + V \end{pmatrix} \end{aligned}$$

Integrating over the flagellum length,  $s \in [0, L]$ , gives

$$-U \begin{pmatrix} C_T L + \epsilon^2 (C_N - C_T) \int_0^L ds h_s^2 \\ \epsilon (C_N - C_T) \int_0^L ds h_s \end{pmatrix} - V \begin{pmatrix} \epsilon (C_N - C_T) \int_0^L ds h_s \\ C_N L \end{pmatrix} - \begin{pmatrix} \epsilon^2 (C_N - C_T) \int_0^L ds h_s h_t \\ \epsilon C_N \int_0^L ds h_t \end{pmatrix}$$

Clearly the term  $\epsilon^2 (C_N - C_T) \int_0^L ds h_s^2$  is a lower order than  $C_T L$  and hence the former is dropped.

The cell body drag follows by setting  $h = 0$  and replacing parameters dependent on geometry with those of the cell body. Hence the cell body drag is

$$-U \begin{pmatrix} C_T^b L_b \\ 0 \end{pmatrix} - V \begin{pmatrix} 0 \\ C_N^b L_b \end{pmatrix}$$

Thus, using Eqn.(7),

$$(C_T^b L_b + C_T L)U = -(C_N - C_T) \left[ \epsilon^2 \int_0^L ds h_s h_t + \epsilon V \int_0^L ds h_s \right],$$

and

$$V = -\frac{1}{C_N^b L_b + C_N L} \left[ \epsilon C_N \int_0^L ds h_t + \epsilon U (C_N - C_T) \int_0^L ds h_s \right].$$

Substituting the expression for  $V$  into the expression for  $U$  we have

$$(C_T^b L_b + C_T L + O(\epsilon^2))U = -\epsilon^2 (C_N - C_T) \left[ \int_0^L ds h_s h_t - \frac{C_N}{C_N^b L_b + C_N L} \int_0^L ds h_s \int_0^L ds h_t \right].$$

We can drop the  $O(\epsilon^2)$  on the left as it is asymptotically small relative to  $C_T L$ .

Hence we have at leading order

$$U = \epsilon^2 \frac{C_T - C_N}{C_N^b L_b + C_N L} \left[ \int_0^L ds h_s h_t - \frac{C_N}{C_N^b L_b + C_N L} \int_0^L ds h_s \int_0^L ds h_t \right]$$

and we recover the expression in the lecture notes provided

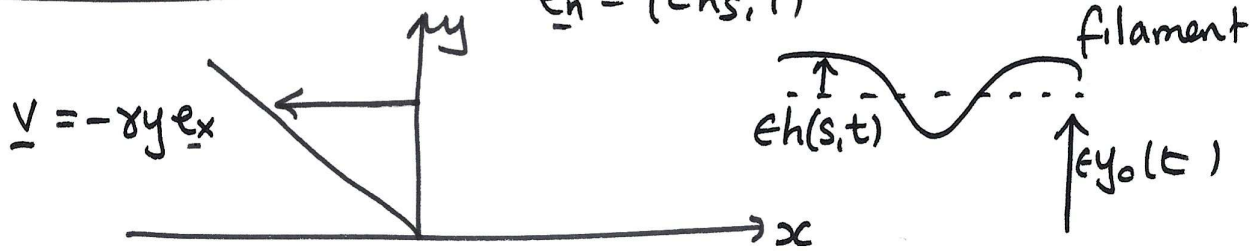
$$\frac{C_N}{C_N^b L_b + C_N L} \left| \frac{\int_0^L ds h_s \int_0^L ds h_t}{\int_0^L ds h_s h_t} \right| \ll 1.$$

# Question 5

$$\underline{e}_t = (-1, \epsilon h_s)$$

$$\underline{e}_n = (\epsilon h_s, 1)$$

displacement  $y_0$  is  $O(\epsilon)$  on assumption there is no drift



$\therefore$  On filament,  $\underline{v} = -\gamma \epsilon (y_0 + h) \underline{e}_x$ ,  $\underline{u} = (u, \epsilon y_{0t} + \epsilon h_t)$

$$\underline{u} - \underline{v} = \begin{pmatrix} u - \gamma \epsilon (y_0 + h) \\ \epsilon y_{0t} + \epsilon h_t \end{pmatrix}$$

$$\boxed{\epsilon y_{0t} \equiv V \text{ in Q4}}$$

$\underline{0}$  = Drag force  $\underline{f}$  = drag per unit length

$$\underline{f} = -(C_N - C_T) \underline{e}_n \cdot (\underline{u} - \underline{v}) \underline{e}_n - C_T (\underline{u} - \underline{v})$$

$$= \underbrace{f}_{\text{Q4}} + (C_N - C_T) \underbrace{(\underline{e}_n \cdot \underline{v}) \underline{e}_n}_{-\gamma \epsilon^2 h_s (y_0 + h) (\epsilon h_s)} + \underbrace{C_T \underline{v}}_{-C_T \gamma \epsilon (y_0 + h)}$$

same expression as in Q4 with

$L_b \equiv 0$  as no cell body and  $V \rightarrow \epsilon y_{0t}$

$$\int_0^L (y_0 + h) ds$$

From zero net drag

$$\underline{0} = -u \left( \frac{C_T L + O(\epsilon^2)}{\epsilon (C_N - C_T) \int_0^L ds h_s} \right) - \left( \frac{\epsilon (C_N - C_T) \int_0^L ds h_s}{C_N L} \right) \epsilon y_{0t}$$

$$- \left( \frac{\epsilon^2 (C_N - C_T) \int_0^L ds h_s h_t}{\epsilon C_N \int_0^L ds h_t} \right) - \left( \frac{O(\epsilon^3)}{\gamma \epsilon^2 (C_N - C_T) \int_0^L ds h_s (y_0 + h)} + \gamma \epsilon C_T \int_0^L (y_0 + h) ds \right)$$



$$\therefore (C_T L + O(\epsilon^2)) U = -\epsilon (C_N - C_T) (\epsilon y_0 t) \int_0^L ds h_s + O(\epsilon^3) \\ - \epsilon^2 (C_N - C_T) \int_0^L ds h_s h_t + \gamma \epsilon C_T \underbrace{\int_0^L (y_0 + h) ds}_{(Ly_0 + \int_0^L h ds)}$$

$$(C_N L) (\epsilon y_0 t) = -\epsilon U (C_N - C_T) \int_0^L ds h_s - \epsilon C_N \int_0^L ds h_t \\ - \gamma \epsilon^2 (C_N - C_T) \int_0^L ds h_s (y_0 + h)$$

Leading order

from  $\epsilon y_0 t$   
↓

$$(C_T L + O(\epsilon^2)) U = -\epsilon \{ O(\epsilon) \} + O(\epsilon^3) + O(\epsilon^2) \\ + \gamma \epsilon C_T L y_0 + \gamma \epsilon C_T \int_0^L h ds$$

$$\epsilon y_0 = \int_0^t d\bar{t} \epsilon y_0 \bar{t} = -\frac{\epsilon}{C_N L} \left[ (C_N - C_T) \int_0^t d\bar{t} \int_0^L ds h_s \right. \\ \left. + C_N \int_0^t d\bar{t} \int_0^L ds h_t \right] + O(\epsilon^2)$$

assuming filament  
starts at  $y_0 \equiv 0$

Aside. No drift entails  $h(s, t)$  is such that these  
integrals remain  $O(1)$  for large time

$$\therefore U = -\frac{\gamma \epsilon}{C_N L} \left[ (C_N - C_T) \int_0^t d\bar{t} \int_0^L ds h_s + C_N \int_0^t d\bar{t} \int_0^L ds h_t \right] \\ + \gamma \epsilon / L \int_0^L h ds + O(\epsilon^2)$$